

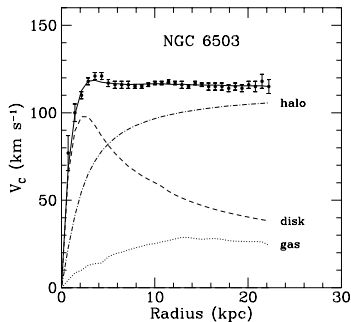
MODEL OF DARK MATTER AND DARK ENERGY BASED ON GRAVITATIONAL POLARIZATION

Luc Blanchet

Gravitation et Cosmologie (GR ϵ CO)
Institut d'Astrophysique de Paris

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Rotation curves of galaxies are flat



For a circular orbit
$$v(r) = \sqrt{\frac{GM(r)}{r}}$$

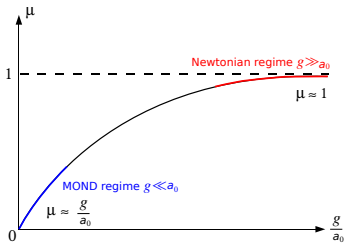
The fact that $v(r)$ is approximately constant implies that beyond the optical disk

$$M_{\text{halo}}(r) \approx r \quad \rho_{\text{halo}}(r) \approx \frac{1}{r^2}$$

Modified Newtonian dynamics (MOND) [Milgrom 1983]

The Newtonian gravitational field is modified in an *ad hoc* way

$$\mu\left(\frac{g}{a_0}\right) \mathbf{g} = \mathbf{g}_{\text{Newtonian}} \quad \text{with} \quad a_0 \approx 1.2 \cdot 10^{-10} \text{ m/s}^2$$



This value of a_0 is mysteriously close to the acceleration scale associated with the cosmological constant Λ

$$a_0 \approx 1.3 a_\Lambda \quad \text{where} \quad a_\Lambda = \frac{1}{2\pi} \left(\frac{\Lambda}{3}\right)^{1/2}$$

Particle dark matter versus MOND

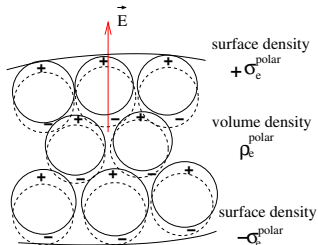
The alternative seems to be

- 1 Either accept the existence of cold dark matter, but
 - made of unknown non-baryonic particles yet to be discovered
 - which fails to reproduce in a natural way the flat rotation curves of galaxies
- 2 Or postulate a modification of the fundamental law of gravity: MOND and its relativistic extensions (TeVeS [Bekenstein 2004] and Einstein-æther), but
 - on an *ad hoc* and not yet physically justified basis

Here we shall adopt a different approach

- 1 Keep the standard law of gravity namely general relativity and its Newtonian limit when $c \rightarrow \infty$
- 2 Use the phenomenology of MOND to guess what could be the (probably unorthodox) nature of dark matter
- 3 Propose a natural physical mechanism for MOND
- 4 Develop a relativistic matter model and apply it to cosmology

The electric field in a dielectric medium



The atoms in a dielectric are modelled by electric dipole moments

$$\boldsymbol{\pi}_e = q \boldsymbol{\xi}$$

The polarization vector is

$$\boldsymbol{\Pi}_e = n \boldsymbol{\pi}_e$$

Density of polarization charges $\rho_e^{\text{polar}} = -\nabla \cdot \boldsymbol{\Pi}_e$

$$\nabla \cdot \mathbf{E} = \frac{\rho_e + \rho_e^{\text{polar}}}{\epsilon_0} \iff \nabla \cdot \left(\mathbf{E} + \frac{\boldsymbol{\Pi}_e}{\epsilon_0} \right) = \frac{\rho_e}{\epsilon_0}$$

The dipoles and polarization vector are aligned with the electric field

$$\frac{\boldsymbol{\Pi}_e}{\epsilon_0} = \underbrace{\chi_e(E)}_{\text{electric susceptibility}} \mathbf{E}$$

Interpretation of the MOND equation [Blanchet 2006]

The MOND equation viewed as a modified Poisson equation [Bekenstein & Milgrom 1984]

$$\nabla \cdot \left[\underbrace{\mu\left(\frac{g}{a_0}\right)}_{\text{MOND function}} \mathbf{g} \right] = -4\pi G \rho$$

is **formally analogous** to the equation of electrostatics inside a dielectric. We pose

$$\mu = 1 + \underbrace{\chi(g)}_{\text{gravitational susceptibility}} \quad \text{and} \quad \underbrace{\mathbf{\Pi}}_{\text{gravitational polarization}} = -\frac{\chi}{4\pi G} \mathbf{g}$$

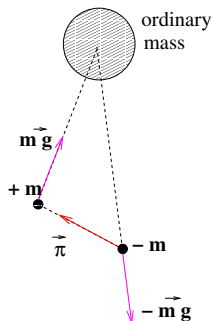
The MOND equation is equivalent to

$$\Delta U = -4\pi G (\rho + \rho_{\text{polar}})$$

In this interpretation the Newtonian law of gravity is not violated but we are postulating a **new form of dark matter consisting of “polarization masses”** with density

$$\rho_{\text{polar}} = -\nabla \cdot \mathbf{\Pi}$$

Sign of the gravitational susceptibility



The “digravitational” medium is modelled by individual dipole moments π

$$\pi = m \xi$$

$$\mathbf{\Pi} = n \pi$$

We suppose that the dipoles consist of particles doublets with

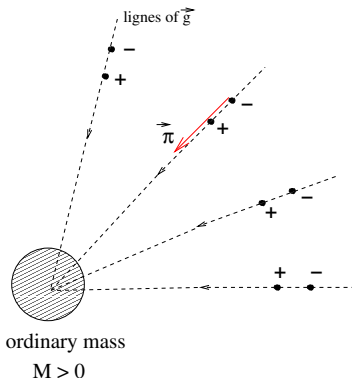
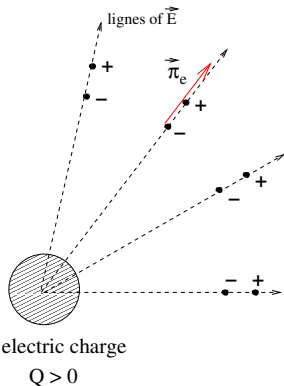
- opposite gravitational masses $m_g = \pm m$
- positive inertial masses $m_i = m$

- 1 The gravitational force is governed by a **negative Coulomb law**
- 2 The dipoles tend to align in the same direction as the gravitational field thus

$$\chi < 0$$

which is nicely compatible with MOND since $0 < \mu < 1 \implies -1 < \chi < 0$

Electric screening versus gravitational anti-screening



Screening by polarization charges

$$\chi_e > 0$$

Anti-screening by polarization masses

$$\chi < 0$$

Interpretation of the dark matter medium

- 1 The dipole moments cannot be stable in a gravitational field and we must invoke some **non-gravitational internal force** (i.e. a “**fifth force**”) which is attractive between unlike masses
- 2 The dark matter medium is like a **gravitational plasma**, i.e. a medium composed of particles

$$(m_i, m_g) = (m, \pm m)$$

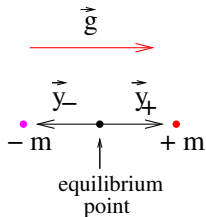
in equal numbers (so the plasma is globally neutral) and with local number densities n_{\pm}

- 3 We find that to obtain MOND we must postulate the Gauss law

$$\nabla \cdot \mathbf{f}_{\text{int}} = -\frac{4\pi G m}{\chi} (n_+ - n_-)$$

where $\chi < 0$ is the susceptibility of the dipolar medium

Gravitational plasma oscillations



Let \mathbf{y}_{\pm} be the displacements from the equilibrium position at which the plasma is locally neutral, *i.e.* $n_{+} = n_{-} = n$. The equation of motion is

$$m \frac{d^2 \mathbf{y}_{\pm}}{dt^2} = \pm m (\mathbf{f}_{\text{int}} + \mathbf{g})$$

Consider a small departure from equilibrium. The density perturbation reads $n_{\pm} = n (1 - \nabla \cdot \mathbf{y}_{\pm})$ to first order in \mathbf{y}_{\pm}

The dipole $\boldsymbol{\xi} = \mathbf{y}_{+} - \mathbf{y}_{-}$ obeys the harmonic oscillator

$$\frac{d^2 \boldsymbol{\xi}}{dt^2} + \omega^2 \boldsymbol{\xi} = 2\mathbf{g}$$

in which ω is the usual **plasma frequency** given here by

$$\omega = \sqrt{-\frac{8\pi G m n}{\chi}}$$

Non viability of the quasi-Newtonian model

The quasi-Newtonian model

- Suggests that the gravitational analogue of the electric polarization is possible
- Yields a simple and natural explanation of the MOND equation
- Requires the existence of a new non-gravitational force
- Interprets the dark matter medium as a polarizable gravitational plasma

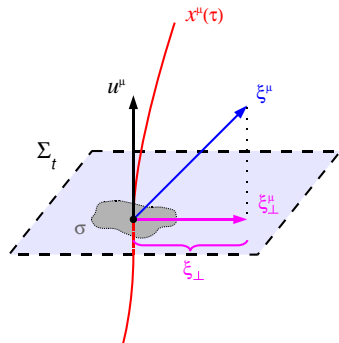
BUT THIS MODEL IS NOT VIABLE

- Is not relativistic
- Involves negative gravitational masses so violates the equivalence principle
- Does not allow to answer questions related to cosmology

We propose a matter action in standard general relativity of the type

$$S = \int d^4x \sqrt{-g} L [J^\mu, \xi^\mu, \dot{\xi}^\mu, g_{\mu\nu}]$$

where the **current density** J^μ and the **dipole moment** ξ^μ are two independent dynamical variables



- The current density $J^\mu = \sigma u^\mu$ is conserved

$$\nabla_\mu J^\mu = 0$$

- The covariant time derivative is denoted

$$\dot{\xi}^\mu \equiv \frac{D\xi^\mu}{d\tau} = u^\nu \nabla_\nu \xi^\mu$$

- Projection perpendicular to the velocity

$$\xi_\perp^\mu \equiv \perp_{\mu\nu} \xi^\nu$$

$$L = \sigma \left[-1 - \sqrt{(u_\mu - \dot{\xi}_\mu)(u^\mu - \dot{\xi}^\mu)} + \frac{1}{2} \dot{\xi}_\mu \dot{\xi}^\mu \right] - \mathcal{W}(\Pi_\perp)$$

- 1 The mass term σ represents the inertial mass density of the dipole moments (say $\sigma = 2m n$)
- 2 The second term is a potential term inspired by the action of particles with spins in general relativity [Papapetrou 1951, Bailey & Israel 1980]
- 3 The third term is a kinetic-like term for the dipole moment and will tell how its evolution differs from parallel transport
- 4 The potential function \mathcal{W} depends on the **polarization field**

$$\Pi_\perp = \sigma \xi_\perp$$

and describes a non-gravitational force \mathcal{F}^μ internal to the dipole moment

Equations of motion and evolution

Varying the action with respect to the dynamical variables J^μ and ξ^μ we find that one equation can be solved with $\left[(u_\mu - \dot{\xi}_\mu)(u^\mu - \dot{\xi}^\mu) \right]^{1/2} = 1$

Equation of motion of dipolar fluid

$$\underbrace{\dot{u}^\mu = -\mathcal{F}^\mu}_{\text{non-geodesic motion}} \quad \text{where} \quad \underbrace{\mathcal{F}^\mu \equiv \hat{\xi}_\perp^\mu \frac{d\mathcal{W}}{d\Pi_\perp}}_{\text{dipolar internal force}}$$

Equation of evolution of dipole moment

$$\dot{\Omega}^\mu = \frac{1}{\sigma} \nabla^\mu \left(\mathcal{W} - \Pi_\perp \frac{d\mathcal{W}}{d\Pi_\perp} \right) - \underbrace{\xi_\perp^\nu R^\mu_{\rho\nu\sigma} u^\rho u^\sigma}_{\text{coupling to Riemann curvature}}$$

$$\text{where} \quad \Omega^\mu \equiv \perp_\nu^\mu \dot{\xi}_\perp^\nu + u^\mu \left(1 + \xi_\perp \frac{d\mathcal{W}}{d\Pi_\perp} \right)$$

The equations depend only on the (space-like) perpendicular projection $\xi_\perp^\mu = \perp_\nu^\mu \xi^\nu$ which represents the **physical dipole moment variable**

Stress-energy tensor

Varying the action with respect to the metric we obtain

$$T^{\mu\nu} = \underbrace{r}_{\text{energy density}} u^\mu u^\nu + \underbrace{\mathcal{P}}_{\text{pressure}} \perp^{\mu\nu} + 2 \underbrace{Q^{(\mu}}_{\text{heat flux}} u^{\nu)} + \underbrace{\Sigma^{\mu\nu}}_{\text{anisotropic stresses}}$$

where

$$\begin{aligned} r &= \rho + \mathcal{W} - \Pi_\perp \mathcal{W}' \\ \mathcal{P} &= -\mathcal{W} + \frac{2}{3} \Pi_\perp \mathcal{W}' \\ Q^\mu &= \sigma \hat{\xi}_\perp^\mu + \Pi_\perp \mathcal{W}' u^\mu - \Pi_\perp^\lambda \nabla_\lambda u^\mu \\ \Sigma^{\mu\nu} &= \left(\frac{1}{3} \perp^{\mu\nu} - \hat{\xi}_\perp^\mu \hat{\xi}_\perp^\nu \right) \Pi_\perp \mathcal{W}' \end{aligned}$$

The mass density ρ contains the rest mass σ and a dipolar term $-\nabla_\lambda \Pi_\perp^\lambda$ which appears as a relativistic generalisation of the polarization mass density

$$\rho = \sigma - \nabla_\lambda \Pi_\perp^\lambda$$

Cosmological perturbations at large scales

- 1 Apply first-order perturbation theory around FLRW background

$$g_{\mu\nu} = \bar{g}_{\mu\nu} + \delta g_{\mu\nu}$$

$$u^\mu = \bar{u}^\mu + \delta u^\mu$$

$$\xi_\perp^\mu = 0 + \delta \xi_\perp^\mu \quad \Leftarrow \text{dipole moment is perturbative}$$

$$\Pi_\perp = 0 + \delta \Pi_\perp \quad \Leftarrow \text{polarization field is perturbative}$$

- 2 Suppose that the potential takes the form

$$\mathcal{W}(\Pi_\perp) = \underbrace{\mathcal{W}_0}_{\text{cosmological constant}} + \frac{1}{2} \mathcal{W}_2 \Pi_\perp^2 + \mathcal{O}(\Pi_\perp^3)$$

- 3 Apply standard **SVT gauge-invariant formalism** with $\delta \xi_\perp^\mu = (0, D^i y + y^i)$

- equations of motion

$$V' + \mathcal{H}V + \Phi = -\mathcal{W}_2 \bar{\sigma} a^2 y$$

$$V_i' + \mathcal{H}V_i = -\mathcal{W}_2 \bar{\sigma} a^2 y_i$$

- evolution equations

$$y'' + \mathcal{H}y' - \mathcal{W}_2 \bar{\sigma} a^2 y = 0$$

$$y_i'' + \mathcal{H}y_i' - \mathcal{W}_2 \bar{\sigma} a^2 y_i = 0$$

Agreement with the Λ -CDM scenario

- ① To first order in the perturbation, $T^{\mu\nu} = T_{\text{de}}^{\mu\nu} + T_{\text{dm}}^{\mu\nu}$ where

$$\begin{aligned} T_{\text{de}}^{\mu\nu} &= -\mathcal{W}_0 g^{\mu\nu} \\ T_{\text{dm}}^{\mu\nu} &= \rho u^\mu u^\nu + \underbrace{2Q^{(\mu} u^{\nu)}}_{\text{heat flux term}} \end{aligned}$$

- ② Posing $\tilde{u}^\mu = u^\mu + Q^\mu/\bar{\rho}$ we can recast the dark matter fluid into a **perturbed pressureless perfect fluid** at linear perturbation order

$$\begin{aligned} T_{\text{dm}}^{\mu\nu} &= \rho \tilde{u}^\mu \tilde{u}^\nu \\ \text{where } \rho &= \underbrace{\sigma}_{\text{rest mass energy density}} - \underbrace{D_i \Pi_\perp^i}_{\text{dipolar polarization energy density}} \end{aligned}$$

The dipolar fluid is undistinguishable from

- **standard DE** (a cosmological constant $\Lambda = 8\pi\mathcal{W}_0$)
- **standard CDM** (a pressureless perfect fluid)

at the level of first-order cosmological perturbations

Non-relativistic limit of the model ($c \rightarrow +\infty$)

- ① Equation of motion of the dipolar particle

$$\frac{d\mathbf{v}}{dt} = \underbrace{\mathbf{g}}_{\text{local gravitational field}} - \underbrace{\mathcal{F}}_{\text{internal force}}$$

- ② Evolution equation of the dipole moment

$$\frac{d^2\xi}{dt^2} = \mathcal{F} + \frac{1}{\sigma} \nabla (\mathcal{W} - \Pi \mathcal{W}') + \underbrace{(\xi \cdot \nabla) \mathbf{g}}_{\text{tidal effect}}$$

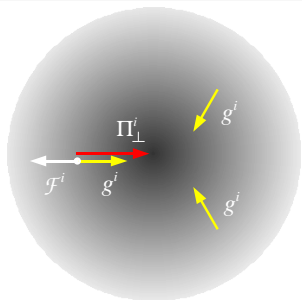
- ③ Conservation of the number of particles

$$\partial_t \sigma = -\nabla \cdot (\sigma \mathbf{v})$$

- ④ Poisson equation for the gravitational potential

$$\nabla \cdot \mathbf{g} = -4\pi \left(\underbrace{\rho_b}_{\text{baryonic matter}} + \sigma - \nabla \cdot \Pi \right)$$

Dipolar dark matter in a central gravitational field



In spherical symmetry we find that there is a solution for which the dipolar fluid is **static (i.e. at rest)**, the dipole moment is **aligned with the gravitational field**, and varies on a very long time-scale so that it is **practically in equilibrium**

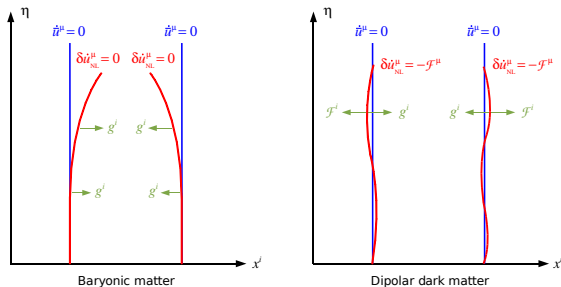
$$\mathbf{v} = \mathbf{0}$$

$$\mathbf{\Pi} = -\Pi(r) \mathbf{e}_r$$

In this solution the **internal force balances the gravitational field**

$$\mathcal{F} = g$$

Weak clustering of dipolar dark matter



- Baryonic matter follows the geodesic equation $\dot{u}^\mu = 0$, therefore collapses in regions of overdensity
- Dipolar dark matter obeys $\dot{u}^\mu = -\mathcal{F}^\mu$, with the internal force \mathcal{F}^i balancing the gravitational field g^i created by an overdensity

We expect that the mass density of dipolar dark matter in a galaxy at low redshift remains close to its mean cosmological value

$$\sigma \approx \bar{\sigma} \ll \rho_b \quad \text{and} \quad v \approx \mathbf{0}$$

Recovering MOND in a galaxy at low red-shift

Using $v \approx 0$ in the equation of motion

$$\mathbf{g} = \hat{\Pi} \frac{d\mathcal{W}}{d\Pi} \quad \leftarrow \text{polarization of the dipolar medium}$$

Using $\sigma \approx \bar{\sigma} \ll \rho_b$ in the gravitational field equation

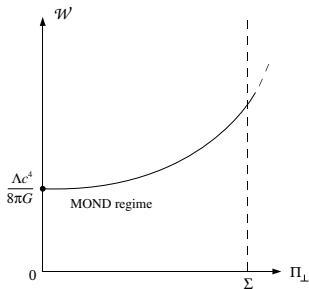
$$\nabla \cdot [\mathbf{g} - 4\pi \Pi] = -4\pi \rho_b \quad \leftarrow \text{modification of the Poisson equation}$$

- 1 Since the polarization Π is aligned with the gravitational field \mathbf{g} we recover the MOND equation with MOND function $\mu = 1 + \chi$ where

$$\Pi = -\frac{\chi}{4\pi} \mathbf{g}$$

- 2 There is one-to-one correspondence between μ and the potential \mathcal{W}

The fundamental potential \mathcal{W}



The potential \mathcal{W} is determined through third order

$$\mathcal{W} = \frac{\Lambda}{8\pi} + 2\pi \Pi_{\perp}^2 + \underbrace{\frac{16\pi^2}{3a_0} \Pi_{\perp}^3}_{\text{MOND acceleration appears here}} + \mathcal{O}(\Pi_{\perp}^4)$$

Though the model is purely *classical* it is tempting to interpret the cosmological constant Λ as a “**vacuum polarization**”, i.e. the residual polarization that remains when the “classical” part of the polarization $\Pi_{\perp} \rightarrow 0$

Order of magnitude of the cosmological constant

- 1 Introduce a purely numerical coefficient α such that

$$a_0 = \frac{1}{2\pi\alpha} \left(\frac{\Lambda}{3} \right)^{1/2} \quad \left(\text{i.e.} \quad a_0 = \frac{a_\Lambda}{\alpha} \right)$$

- 2 Write the fundamental potential as $\mathcal{W} = \frac{3\pi a_0^2}{2} f\left(\frac{\Pi_\perp}{a_0}\right)$ with

$$f(x) = \underbrace{\alpha^2 + \frac{4}{3}x^2 + \frac{32\pi}{9}x^3 + \mathcal{O}(x^4)}_{\text{some "universal" function of } x \equiv \Pi_\perp/a_0}$$

- 3 The numerical coefficients in $f(x)$ are expected to be of the order of one, hence the cosmological constant should be of the order of

$$\Lambda \sim a_0^2$$

in good agreement with the observations (which give $\alpha \approx 0.8$)

Conclusions

This model based on gravitational polarization:

- 1 Offers a natural physical explanation of the phenomenology of MOND
- 2 Benefits from both the successes of MOND at galactic scales and of Λ -CDM at cosmological scales
- 3 Yields a nice unification between the dark energy in the form of Λ and the dark matter *à la* MOND
- 4 But is “phenomenological” and presumably only valid in a regime of weak gravitational fields

More work should be done to:

- 1 Connect the model to fundamental (quantum) physics valid at microscopic scale [work in progress]
- 2 Investigate second-order perturbations in cosmology
- 3 Numerically compute the non-linear growth of perturbations and compare with large-scale structures
- 4 Test the model at the intermediate scale of galaxy clusters